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Electroweak symmetry breaking and singlino dark matter with deflected anomaly mediation

Norimi Yokozaki

Graduate School of Science, Hiroshima University, Higashi-Hiroshima, Hiroshima 739-8526, Japan

E-mail: norimi@theo.phys.sci.hiroshima-u.ac.jp

ABSTRACT: We investigate the phenomenology of the Nearly Minimal Supersymmetric Standard Model (nMSSM) in the deflected anomaly mediation scenario. We also include the Fayet-Iliopoulos D-term of the standard model gauge group. In the nMSSM, the mu term is replaced by the vacuum expectation value of the gauge singlet; therefore, there is no difficulty in generating the B-term of the SUSY breaking scale. Although the messenger sector is introduced, direct couplings between nMSSM fields and messenger sector fields are forbidden by the discrete symmetry. Therefore, the phenomenology at the weak scale does not depend on the detail of the messenger sector. We show that there are regions of parameter space in which electroweak symmetry breaking occurs successfully and the lightest Higgs is heavier than the LEP bound. We show that the gluino is light in this scenario. The lightest neutralino, which is mainly composed of a singlino, is a candidate for dark matter. The relic density explains the observed abundance of dark matter. The dark matter-nucleon scattering cross section satisfies the current limit from CDMS and XENON10 with a small value for the strange quark content of a nucleon.

KEYWORDS: Supersymmetry Breaking, Supergravity Models

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1 Introduction

The Minimal Supersymmetric Standard Model (MSSM) is the most attractive framework for the physics beyond the Standard Model. In the MSSM, gauge coupling unification is achieved and the Higgs potential is stabilized. Despite these good features, the MSSM has difficulty in the Higgs sector. The MSSM has a μ term, $\mu H_1 H_2$, in the superpotential. To maintain the weak-scale vacuum expectation value (VEV) of a Higgs, $|\mu|$ has to be at the weak scale. However, it is difficult to explain why such a dimensionful parameter is much smaller than Plank scale or GUT scale. This problem is the so-called μ problem.

A simple way of solving the μ problem is to introduce a gauge singlet, and replace the μ by the VEV of the gauge singlet field:

$$\mu H_1 H_2 \to \lambda \langle S \rangle H_1 \cdot H_2 . \tag{1.1}$$

The most famous model to include a gauge singlet is the Next-to-Minimal Supersymmetric Standard Model (NMSSM). In the NMSSM, a new discrete symmetry, Z_3 , is introduced to

forbid the mass term for S. However, Z_3 symmetry spontaneously breaks down when electroweak symmetry breaking occurs. At this point, unacceptably large cosmological domain walls appear [1]. In the Nearly Minimal Supersymmetric Standard Model (nMSSM) [2–4], the cosmological domain wall problem is solved by tadpoles. The tadpoles are generated by supergravity interactions and explicitly break the discrete symmetry. Therefore, the domain wall problem does not arise. The nMSSM has the same attractive feature of electroweak baryogenesis as the NMSSM has. To achieve successful electroweak baryogenesis, a strong first-order phase transition is required. Therefore, new sources of CP-violation beyond the CKM matrix have to exist. In the nMSSM, there are additional sources of CPviolation in the singlet sector. Therefore, unlike MSSM, nMSSM does not rely on radiative contributions from a light stop for strong first-order phase transition [5, 6].

SUSY breaking terms are important in discussing phenomenology, SUSY breaking effects are transmitted to the nMSSM sector from a hidden sector by one or more mediation schemes. One interesting mediation scheme is anomaly mediation [7–9]. In anomaly mediation, the supergravity actions of the hidden sector and visible sector are sequestered. SUSY breaking effects are transmitted to the visible sector due to the superconformal anomaly. There are studies in which soft breaking terms are derived by anomaly mediation with NMSSM-like models [10, 11]. In these works, successful electroweak symmetry-breaking is achieved; however, the VEV of S is on the order of a few TeV. This leads to a large higgsino mass: $\mu_{\text{eff}} = \lambda \langle S \rangle$. Therefore, there are large mass splittings among Higgs (and neutralinos).

To obtain a moderate value for μ_{eff} , we consider a deflected anomaly mediation scenario [12–14], which introduces an additional messenger sector. The SUSY breaking mass for a messenger is given by a VEV of the gauge singlet field, X. In the original deflected anomaly mediation scenario [12, 13], the superpotential of X is extremely flat; therefore, the fermionic component of X, ψ_X , becomes light and the lightest SUSY particle is ψ_X . In the positively deflected anomaly mediation scenario [14], the superpotential is not flat; therefore, ψ_X does not have to be light [15] and an ordinary SUSY particle can be a candidate for dark matter. We consider the positively deflected anomaly mediation scenario. We also consider SUSY breaking with the Fayet-Iliopoulos D-term.

We show that when nMSSM and deflected anomaly mediation are combined, successful electroweak symmetry breaking occurs for a moderate value of μ_{eff} . We also show that the lightest neutralino, which is mainly composed of a singlino, is a good candidate for dark matter. We also present sparticle mass spectra.

This paper is organized as follows. In section 2, we introduce the nMSSM Lagrangian and discuss tadpoles. We also discuss the direct couplings between nMSSM fields and messenger sector fields. In section 3, we derive the soft SUSY breaking terms of the nMSSM fields in the deflected anomaly mediation scenario. Section 4 is devoted to the phenomenology of this scenario. Finally, section 5 presents our conclusions.

2 Nearly minimal supersymmetric standard model

In this section, we discuss tadpoles and direct couplings between nMSSM fields and messenger sector fields. First, we introduce the nMSSM Lagrangian. The superpotential and soft breaking terms in the nMSSM are

$$W_{\rm nMSSM} = \lambda \hat{S} \hat{H}_1 \cdot \hat{H}_2 + \frac{m_{12}^2}{\lambda} \hat{S} + y_u \hat{Q} \cdot \hat{H}_2 \hat{U}^c + y_d \hat{Q} \cdot \hat{H}_1 \hat{D}^c + y_l \hat{L} \cdot \hat{H}_1 \hat{E}^c , \qquad (2.1)$$

and

$$-\mathcal{L}_{\text{soft}} = m_S^2 |S|^2 + (a_\lambda S H_1 \cdot H_2 + h.c.) + (t_S S + h.c.) + \tilde{m}_{H_1}^2 H_1^{\dagger} H_1 + \tilde{m}_{H_2}^2 H_2^{\dagger} H_2 + \tilde{m}_Q^2 \tilde{Q}^{\dagger} \tilde{Q} + \tilde{m}_U^2 |\tilde{u}_R|^2 + \tilde{m}_D^2 |\tilde{d}_R|^2 + \tilde{m}_L^2 \tilde{L}^{\dagger} \tilde{L} + \tilde{m}_E^2 |\tilde{e}_R|^2 + (a_u \tilde{Q} \cdot H_2 \tilde{u}_R^* + a_d \tilde{Q} \cdot H_1 \tilde{d}_R^* + a_l \tilde{L} \cdot H_1 \tilde{e}_R^* + h.c.) .$$
(2.2)

 \hat{S} denotes a gauge singlet chiral superfield, and S is the scalar component of \hat{S} . When S acquires the VEV, the higgsino mass parameter, $\mu_{\text{eff}} = \lambda v_s$, is generated effectively. We take λ to be real positive by suitable redefinitions of S, H_1 and H_2 . Unlike the NMSSM, there are no trilinear terms for the gauge singlet. $m_{12}^2 \hat{S}/\lambda$ and $t_S S$ are tadpoles. They are absent at the tree level; however, they are generated radiatively by supergravity interactions. These terms are on the order of the weak scale, as we describe below.

2.1 Tadpoles

The greatest difference between the nMSSM and NMSSM is the existence of tadpoles in the former. The tadpoles are generated by supergravity interaction. In the nMSSM, the theory has a global discrete symmetry at tree level. This symmetry guarantees that the generated tadpoles are on the order of the weak scale, despite the fact that supergravity interactions break global symmetries. Because the tadpoles explicitly break the discrete symmetry, the domain wall problem does not appears.

In nMSSM, the Lagrangian has a discrete R symmetry $Z_{nR'}$ at tree level. The charge assignment of the fields is shown in table 1. The charge of $Z_{nR'}$, $Q_{nR'}$, is defined as

$$Q_{PQ} + 3Q_R, \qquad (2.3)$$

where Q_{PQ} denotes the charge of Peccei-Quinn symmetry, $U(1)_{PQ}$, and Q_R denotes the charge of $U(1)_R$. Under $Z_{nR'}$, the nMSSM fields transform as

$$\Phi_i \to \Phi_i \exp\left(i\frac{Q_{nR'}}{n}\theta\right),$$
(2.4)

where Φ_i denotes nMSSM fields. If the Lagrangian respects $Z_{5R'}$ or $Z_{7R'}$ at the tree level, the scale of the generated tadpoles can naturally be the weak scale [2, 4]. When the discrete symmetry is $Z_{5R'}$, tadpoles of six-loop order are generated. When the discrete symmetry is $Z_{7R'}$, tadpoles of seven-loop order are generated. The tadpoles break the $Z_{5R'}$ or $Z_{7R'}$, and therefore no cosmological domain wall problem exists.

The generated tadpoles are given by [4]:

$$V_{\text{tad}} \sim \frac{1}{(16\pi^2)^l} \left(M_p M_{\text{susy}}^2 S + M_{\text{susy}} F_S + h.c. \right),$$
 (2.5)

	\hat{H}_1	\hat{H}_2	\hat{S}	\hat{Q}	Ĺ	\hat{U}^c	\hat{D}^c	\hat{E}^c	W
$\mathrm{U}(1)_{PQ}$	1	1	-2	-1	-1	0	0	0	0
$\mathrm{U}(1)_R$	0	0	2	1	1	1	1	1	2
$Z_{nR'}$	1	1	4	2	2	3	3	3	6

 Table 1. Charge assignments of fields.

	X	Ψ	$\bar{\Psi}$	$W_{\rm mess}$
$\mathrm{U}(1)_{PQ}$	0	0	0	0
$\mathrm{U}(1)_R$	1	1/2	1/2	2
$Z_{nR'}$	3	3/2	3/2	6

Table 2. Charge assignment of the fields in the messenger sector.

where l is the number of the loops at which tadpoles first appear. l = 6 in the $Z_{5R'}$ case and l = 7 in $Z_{7R'}$ case. In a deflected anomaly mediation scenario as well as an anomaly mediation scenario, M_{susy} is $\mathcal{O}(10 \text{TeV})$; therefore, $Z_{7R'}$ is favorable. As we describe below, $Z_{7R'}$ also forbids direct couplings between nMSSM fields and messenger sector fields.

2.2 Direct couplings to the messenger sector

In a deflected anomaly mediation scenario, the messenger sector is introduced in addition to the hidden sector, which is the origin of SUSY breaking. The messenger sector contains a gauge singlet chiral superfield and messenger superfields. The messengers transmit the SUSY breaking to the nMSSM sector, and this SUSY breaking is comparable to that of anomaly mediation. In this subsection, we show that direct couplings between the messenger sector fields and the nMSSM fields do not exist.

We consider the following superpotential in the messenger sector.

$$W_{\rm mess} = \frac{1}{2}m_X \hat{X}^2 + \lambda_X \hat{X} \bar{\Psi}_i \Psi^i \,, \qquad (2.6)$$

where \hat{X} is a gauge singlet chiral superfield. $\bar{\Psi}_i$ and Ψ^i are the messenger fields that transform $\bar{\mathbf{5}}$ and $\mathbf{5}$ for the SU(5) GUT gauge group respectively. The $Z_{nR'}$ charge assignment of the fields is shown in table 2.

In this charge assignment, there are no direct couplings between messenger sector fields and nMSSM fields. A direct coupling between the messengers and nMSSM gauge singlet, $S\bar{\Psi}_i\Psi^i$, is forbidden by $Z_{nR'}$ symmetry. SX^2 and XH_uH_d terms are also forbidden. On the other hand, the S^2X term is forbidden by $Z_{7R'}$ but allowed by $Z_{5R'}$. In the discussion about tadpoles in the previous subsection, we assumed that $Z_{7R'}$ symmetry exists at tree level. Therefore, there are no direct couplings between nMSSM fields and the messenger sector fields. The phenomenology of the nMSSM at the weak scale does not depend on the detail of messenger sector.

3 SUSY breaking

In this section, we derive the SUSY breaking terms of the nMSSM in the deflected anomaly mediation scenario. We also show the corrections to the soft scalar mass with the Fayet-Iliopoulos D-term.

In the original deflected anomaly mediation scenario, the superpotential of the gauge singlet \hat{X} is flat; therefore in general, the lightest SUSY particle (LSP) is the fermionic component of X, ψ_X . Threshold corrections to the sparticle mass squared are negative. In the positively deflected anomaly mediation scenario, the superpotential of \hat{X} is not flat; therefore, ψ_X is not necessarily the LSP. Corrections to the sparticle mass squared are positive. We consider the positively deflected anomaly mediation scenario. In appendix A, we give an explicit example of the positively deflected anomaly mediation scenario in which the fermionic partner of X is not the LSP.

When \hat{X} acquires the VEV, messengers obtain SUSY breaking mass, $X + F_X \theta^2$. This SUSY breaking mass introduces an intermediate threshold that depends on θ^2 . In deflected anomaly mediation, corrections from anomaly mediation to soft breaking terms are generated by the following threshold.

$$\frac{X + F_X \theta^2}{\Lambda \hat{\phi}} = \frac{X}{\Lambda} \left[1 + \left(\frac{F_X}{X} - F_\phi \right) \theta^2 \right] \equiv \frac{X}{\Lambda} \left(1 + dF_\phi \theta^2 \right), \tag{3.1}$$

where $\hat{\phi}$ is the chiral compensator field, $\hat{\phi} = 1 + F_{\phi}\theta^2$, and Λ is the ultraviolet cutoff. d is the deflection parameter, which denotes the threshold correction to the SUSY breaking. In the positively deflected anomaly mediation scenario, d is positive.

Messengers also affect the beta-functions of gauge couplings. The beta-functions of gauge couplings above the scale |X| are written as

$$\frac{dg_a}{d\ln\mu} = -\frac{g_a^3}{16\pi^2}(b_a - N_f),$$
(3.2)

where N_f is the number of messengers. For the intermediate threshold and modification of the beta-functions, soft breaking terms in an anomaly mediation are changed to those of a deflected anomaly mediation scenario.

In a deflected anomaly mediation, the gaugino mass, soft breaking mass and scalar trilinear couplings at the scale μ are obtained using the following relations [12, 14].

$$\frac{m_{\lambda}(\mu)}{g^{2}(\mu)} = -\frac{F_{\phi}}{2} \left(\frac{\partial}{\partial \ln \mu} - d \frac{\partial}{\partial \ln |X|} \right) g^{-2} \left(\frac{\mu}{\Lambda}, \frac{|X|}{\Lambda} \right)
\tilde{m}_{i}^{2}(\mu) = -\frac{|F_{\phi}^{2}|}{4} \left(\frac{\partial}{\partial \ln \mu} - d \frac{\partial}{\partial \ln |X|} \right)^{2} \ln Z_{i} \left(\frac{\mu}{\Lambda}, \frac{|X|}{\Lambda} \right)
\frac{a_{ijk}(\mu)}{y_{ijk}(\mu)} = -\frac{F_{\phi}}{2} \left(\frac{\partial}{\partial \ln \mu} - d \frac{\partial}{\partial \ln |X|} \right) \sum_{l=i,j,k} \ln Z_{l} \left(\frac{\mu}{\Lambda}, \frac{|X|}{\Lambda} \right),$$
(3.3)

where F_{ϕ} is the F-term of the chiral compensator fields and corresponds to the gravitino mass. |X| is the messenger scale and $\mu < |X|$. d is defined in eq. (3.1). |X| and d can

be determined by the superpotential and the soft breaking terms in the messenger sector (see appendix A). However, we treat them as the parameters of SUSY breaking because we focus on the phenomenology at the weak scale.

In general, the formula for soft breaking terms is complicated. However, by setting the scale as $\mu = |X|$, the soft breaking terms are simplified as

$$m_{\lambda_{a}} = -\frac{g_{a}^{2}}{(4\pi)^{2}} (b_{a} - dN_{f}) F_{\phi},$$

$$\tilde{m}_{i}^{2} = \frac{|F_{\phi}|^{2}}{2(4\pi)^{4}} \sum_{a} c_{a}^{i} g_{a}^{4} [b_{a} + d(d+2)N_{f}], -\frac{|F_{\phi}|^{2}}{4} \sum_{y} \frac{\partial \gamma_{i}(|X|)}{\partial y} \beta_{y}(|X|),$$

$$a_{ijk} = -\frac{F_{\phi}}{2} [\gamma_{i}(|X|) + \gamma_{j}(|X|) + \gamma_{k}(|X|)] y_{ijk}.$$
(3.4)

Here, N_f is the number of messengers. b_a and c_a^i are the coefficients of the gauge coupling beta functions and the anomalous dimensions of the fields respectively. $b_a = (-33/5, -1, 3)$, $c_a^L = (3/5, 3, 0)$, $c_a^{E^c} = (12/5, 0, 0)$, $c_a^Q = (1/15, 3, 16/3)$, $c_a^{U^c} = (16/15, 0, 16/3)$ and $c_a^{D^c} = (4/15, 0, 16/3)$.

The formula for gaugino masses is easily obtained with

$$g_a^{-2}(\mu) = g_a^{-2}(\Lambda) + \frac{b_a}{8\pi^2} \ln \frac{\mu}{|X|} + \frac{b_a - N_f}{8\pi^2} \ln \frac{|X|}{\Lambda}.$$
(3.5)

Equation (3.5) can be obtained by integrating the beta-functions explicitly. The derivations of \tilde{m}_i^2 and a_{ijk} are given in appendix B.

For the first and second generations of squarks and sleptons, we can neglect the contributions from Yukawa couplings. However, for the soft scalar mass and the A-term of the third generation of squarks and sleptons, the contributions from Yukawa couplings are important. For \tilde{m}_S^2 , $\tilde{m}_{H_1}^2$, $\tilde{m}_{H_2}^2$ and a_{λ} , contributions from Yukawa couplings are also important. The anomalous dimensions of H_1 and H_2 are different from those of the MSSM due to λ and are given in appendix C. The anomalous dimensions of the other fields are same as those of the MSSM and are given in [16]. When Yukawa couplings are small, we obtain the results of [12, 14].

In a supersymmetric model, there is an additional source of SUSY breaking, the Fayet-Iliopoulos D-term. This term contributes to the square of the scalar mass.

The Fayet-Iliopoulos D-term is

$$\mathcal{L} \ni -\xi D. \tag{3.6}$$

The D-term of the Lagrangian is written as

$$\mathcal{L}_D = \frac{1}{2}D^2 - gD\sum_i q_i A_i^{\dagger} A_i - \xi D, \qquad (3.7)$$

where q_i is the U(1) charge of the field A_i . After eliminating the D-term with the equation of motion, the Lagrangian \mathcal{L}_D becomes

$$\mathcal{L}_D = -\frac{1}{2} \left(\sum_i q_i A_i^{\dagger} A_i + \xi \right)^2 \,. \tag{3.8}$$

This leads to additional contributions to the scalar mass terms:

$$\tilde{m}_{ij}^2 \to \tilde{m}_{ij}^2 + q_i \xi \delta_{ij}. \tag{3.9}$$

In the Supersymmetric Standard Model, there is only one U(1) gauge group. In the lepton sector, the hypercharge of the SU(2) doublet is -1 and the hypercharge of the SU(2) singlet is +2. Therefore, we can not solve the tachyonic slepton mass problem in anomaly mediation using only the U(1)_Y D-term.

So far, the additional contributions from the D-term to the soft breaking mass of the nMSSM matter fields are

$$\begin{split} \delta \tilde{m}_{L}^{2} &= -D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \\ \delta \tilde{m}_{E}^{2} &= 2D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \\ \delta \tilde{m}_{Q}^{2} &= \frac{1}{3} D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \\ \delta \tilde{m}_{U}^{2} &= -\frac{4}{3} D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \\ \delta \tilde{m}_{D}^{2} &= \frac{2}{3} D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \\ \delta \tilde{m}_{H_{1}}^{2} &= -D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \\ \delta \tilde{m}_{H_{2}}^{2} &= D_{Y} \frac{|F_{\phi}|^{2}}{(4\pi^{4})}, \end{split}$$
(3.10)

where D_Y comes from the common parameter ξ in eq. (3.7).

We can now evaluate the soft breaking terms of the nMSSM at the messenger scale using eq. (3.4). We solve the renormalization group equations (RGE) using them as the boundary conditions, and then evaluate the soft breaking terms at the weak scale. We use RGE codes contained in the NMSSMTools software package [17, 18]. We also add the D-term contributions in eq. (3.7) to the soft breaking mass. In the next section, we discuss the phenomenology of the nMSSM with the soft breaking terms obtained by deflected anomaly mediation.

4 Phenomenology

In this section, we investigate the phenomenological aspects of the nMSSM. First, we discuss the existence of Landau poles. We demand that λ should not meet the Landau pole up to the scale at which the tadpoles are generated. Next we study the regions of parameter space where successful electroweak breaking occurs, and we evaluate the mass of the lightest Higgs. Subsequently, we discuss the lightest neutralino as a dark matter candidate. We evaluate the relic density of the lightest neutralino. We also discuss the direct detection of dark matter. Finally, we obtain sparticle mass spectra.



Figure 1. The region that is consistent with the perturbativity of λ up to the GUT scale is shown. The gray shaded region below the solid line is allowed. The region above the solid line is excluded owing to the existence of the Landau pole below the GUT scale.

4.1 The Landau pole

In the nMSSM, tadpoles generated by supergravity interaction are proportional to powers of λ [4]. Therefore to maintain the tadpoles at the weak scale, λ should not meet the Landau pole up to the scale, at which the tadpoles are generated. We investigate the region of λ and tan β that satisfies the perturbativity condition below the GUT scale.

The beta-functions of λ and y_t are

$$\beta_{\lambda} = \frac{1}{16\pi^2} \left(4\lambda^2 + 3y_t^2 + 3y_b^2 + y_\tau^2 - \frac{3}{5}g_1^2 - 3g_2^2 \right) \lambda ,$$

$$\beta_{y_t} = \frac{1}{16\pi^2} \left(\lambda^2 + 6y_t^2 + y_b^2 - \frac{13}{15}g_1^2 - 3g_2^2 - \frac{16}{3}g_3^2 \right) y_t .$$
(4.1)

These beta-functions strongly depend on λ and $\tan \beta$ through the top Yukawa coupling. Figure 1 shows the allowed region where the perturbativity is satisfied up to the GUT scale. The calculation is performed using the RGE code included in the NMSSMTools package. The current experimental value of the top mass is 173.1 ± 1.3 GeV [19]. We take the central value for $m_{\rm top}$ as 173.1 GeV. The shaded region is consistent with the perturbativity of λ . The result depends on the value of $m_{\rm top}$ and supersymmetric threshold corrections of α_s . Therefore there is small difference among the results of [20] and [5] and our results. In our calculation, the region where $\tan \beta \gtrsim 2.0$ and $\lambda \lesssim 0.7$ is allowed.

4.2 Electroweak symmetry breaking

In this subsection, we consider the conditions for electroweak symmetry breaking and evaluate $\mu_{\text{eff}} = \lambda \langle S \rangle$. We also evaluate the mass of the lightest Higgs.

After obtaining the soft breaking terms at the weak scale, we now evaluate the Higgs potential, $V = V_{\text{tree}} + \Delta V$. From eqs. (2.1) and (2.2), the tree-level Higgs potential is

written as

$$V_{\text{tree}} = \tilde{m}_{H_1}^2 H_1^{\dagger} H_1 + \tilde{m}_{H_2}^2 H_2^{\dagger} H_2 + m_s^2 |S|^2 + m_{12}^2 (H_1 \cdot H_2 + h.c.) + \lambda^2 |H_1 \cdot H_2|^2 + \lambda^2 |S|^2 (H_1^{\dagger} H_1 + H_2^{\dagger} H_2) + \frac{g^2}{2} |H_1^{\dagger} H_2|^2 + \frac{\bar{g}^2}{8} (H_2^{\dagger} H_2 - H_1^{\dagger} H_1)^2 + (t_s S + h.c.) + (a_\lambda S H_1 \cdot H_2 + h.c.), \qquad (4.2)$$

where $\bar{g}^2 = g^2 + g'^2$. ΔV is the one-loop contribution to the effective potential [21]:

$$\Delta V = \frac{1}{64\pi^2} \left(\sum_b g_b m_b^4 \left[\ln\left(\frac{m_b^2}{Q^2}\right) - \frac{3}{2} \right] - \sum_f g_f m_f^4 \left[\ln\left(\frac{m_f^2}{Q^2}\right) - \frac{3}{2} \right] \right) .$$
(4.3)

 g_b and g_f are the degrees of freedom for bosons and fermions respectively. We determine $\mu_{\text{eff}} \equiv \lambda \langle S \rangle$, t_s and m_{12}^2 using the stationary conditions of the Higgs potential. From eqs. (4.2) and (4.3), the stationary conditions are

$$\begin{aligned} \frac{\partial V}{\partial v_1} &= 2v_1 \left[\tilde{m}_{H_1}^2 + (m_{12}^2 + a_\lambda v_s) \frac{v_2}{v_1} - \frac{\bar{g}^2}{4} (v_2^2 - v_1^2) + \lambda^2 (v_2^2 + v_s^2) + \frac{1}{2v_1} \frac{\partial \Delta V}{\partial v_1} \right] = 0, \\ \frac{\partial V}{\partial v_2} &= 2v_2 \left[\tilde{m}_{H_2}^2 + (m_{12}^2 + a_\lambda v_s) \frac{v_1}{v_2} + \frac{\bar{g}^2}{4} (v_2^2 - v_1^2) + \lambda^2 (v_1^2 + v_s^2) + \frac{1}{2v_2} \frac{\partial \Delta V}{\partial v_2} \right] = 0, \\ \frac{\partial V}{\partial v_s} &= 2v_s \left[m_s^2 + \lambda^2 (v_1^2 + v_2^2) + \frac{t_s}{v_s} + a_\lambda \frac{v_1 v_2}{v_s} + \frac{1}{2v_s} \frac{\partial \Delta V}{\partial v_s} \right] = 0, \end{aligned}$$
(4.4)

where $v_1 = \langle H_1^0 \rangle$, $v_2 = \langle H_2^0 \rangle$ and $v_s = \langle S \rangle$. As we describe later, there is only a small region of parameter space where successful electroweak symmetry breaking occurs with $v_s > 0$; therefore, we take $v_s < 0$. From eq. (4.4), μ_{eff} can be determined by,

$$\mu_{\text{eff}}^{2} = -\frac{M_{Z}^{2}}{2} + \frac{\tilde{m}_{H_{1}}^{2} + \frac{1}{2v_{1}}\frac{\partial\Delta V}{\partial v_{1}} - \left(\tilde{m}_{H_{2}}^{2} + \frac{1}{2v_{2}}\frac{\partial\Delta V}{\partial v_{2}}\right)\tan^{2}\beta}{\tan^{2}\beta - 1}.$$
(4.5)

 μ_{eff} , m_{12}^2 and t_s are determined from eq. (4.4). We now evaluate the Higgs mass. We expand H_1^0 , H_2^0 and S as

$$H_1^0 = v_1 + \frac{1}{\sqrt{2}} \left(h_1^0 + ia_1 \right),$$

$$H_2^0 = v_2 + \frac{1}{\sqrt{2}} \left(h_2^0 + ia_2 \right),$$

$$S^0 = v_s + \frac{1}{\sqrt{2}} \left(s + ia_s \right).$$
(4.6)

Using these expanded fields, the CP-even Higgs mass matrix is written as

$$(h_1^0 h_2^0 S) M^2 \begin{pmatrix} h_1^0 \\ h_2^0 \\ S \end{pmatrix}.$$
 (4.7)

At tree level, the components of M^2 are

$$M_{11}^{2} = s_{\beta}^{2} M_{a}^{2} + c_{\beta}^{2} M_{Z}^{2},$$

$$M_{12}^{2} = -s_{\beta} c_{\beta} \left(M_{a}^{2} + M_{Z}^{2} - 2\lambda^{2} v^{2} \right),$$

$$M_{13}^{2} = v \left(s_{\beta} a_{\lambda} + 2c_{\beta} \lambda^{2} v_{s}^{2} \right),$$

$$M_{22}^{2} = c_{\beta}^{2} M_{a}^{2} + s_{\beta}^{2} M_{Z}^{2},$$

$$M_{23}^{2} = v \left(c_{\beta} a_{\lambda} 2 + s_{\beta} \lambda^{2} v_{s} \right),$$

$$M_{33}^{2} = -\frac{1}{v_{s}} \left(t_{s} + s_{\beta} c_{\beta} a_{\lambda} v_{s} \right),$$
(4.8)

where $c_{\beta} = \cos \beta$ and $s_{\beta} = \sin \beta$. The CP-odd Higgs mass matrix at tree-level is

$$(A^0 a_s) \begin{bmatrix} M_a^2 & -a_\lambda v_s \\ -a_\lambda v_s & -\frac{1}{v_s} \left(t_s + s_\beta c_\beta a_\lambda v^2 \right) \end{bmatrix} \begin{pmatrix} A^0 \\ a_s \end{pmatrix},$$

$$(4.9)$$

where $M_a^2 = -(m_{12}^2 + a_\lambda v_s)/c_\beta s_\beta$. $A^0 = a_d s_\beta + a_u c_\beta$, and its orthogonal combination is absorbed by the Z boson.

We now present the results of numerical calculations. Figure 2 shows the allowed region of successful electroweak symmetry breaking without tachyonic sleptons. We set the messenger scale to $5F_{\phi} \simeq 150$ TeV. Successful electroweak symmetry breaking occurs in the region covered by red squares. In the region covered by blue crosses, the mass of the lightest Higgs satisfies the LEP bound with the electroweak symmetry breaking. When the number of the messengers, N_f increases, the allowed region of the deflection parameter d is shifted downward. Therefore, in the scenario with small d (d < 1), two or more messengers have to exist. For simplicity, we assume that there is one messenger in the following analysis. Although there is a region where successful electroweak symmetry breaking occurs with large tan β and $v_s > 0$, the region is very small. Therefore we take $v_s < 0$.

Figure 3 shows the dependence of μ_{eff} on SUSY breaking. We see that $|\mu_{\text{eff}}|$ is a decreasing function of D_Y , while it is an increasing function of d. This can be understood from eqs. (3.4), (3.10) and (4.5). When D_Y increases, $m_{H_1}^2$ decreases and $m_{H_2}^2$ increases. This implies that $|\mu_{\text{eff}}|$ decreases as D_Y increases. When d increases, $m_{H_1}^2$ and $m_{H_2}^2$ increase at almost the same rate. This implies that $|\mu_{\text{eff}}|$ increases as d increases. In this scenario, moderate values of μ_{eff} , $100 < |\mu_{\text{eff}}| < 550$, are obtained without meeting the Landau pole.

Figure 4 shows the dependence of the lightest Higgs mass on d and D_Y . The calculation is performed with NMSSMTools, including two-loop corrections. We extend the codes to include tadpoles. In this scenario, the mass of the lightest Higgs can be heavier than the LEP bound.

4.3 Dark matter

In this scenario, the lightest neutralino is the LSP in the wide range of parameter space. Therefore, the lightest neutralino is a candidate for dark matter. In this subsection, we evaluate the relic density of the lightest neutralino, which is mainly composed of a singlino. We also calculate the the neutralino-proton scattering cross section, and discuss the direct detection of dark matter.



Figure 2. Successful electroweak symmetry breaking occurs in the region covered by red squares, and the region covered by blue crosses satisfies the Higgs mass bound of the LEP. In other regions, the sleptons are tachyonic. The calculation is performed with $\lambda = 0.69$ and $m_0 = F_{\phi}/(4\pi)^4 =$ 200 GeV. The messenger scale is taken to be $5F_{\phi}$. $\tan\beta$ and the number of messengers N_f are $\tan\beta = 2$ and $N_f = 1$ in the top-left figure, $\tan\beta = 3$ and $N_f = 1$ in the top-right figure, $\tan\beta = 2$ and $N_f = 2$ in the bottom-left figure and $\tan\beta = 20$ and $N_f = 1$ in the bottom-right figure. The bottom-right figure is evaluated with $v_s > 0$. The others are evaluated with $v_s < 0$.



Figure 3. The values of $|\lambda v_s|$ are shown. The calculations are performed with $\lambda = 0.69$, $\tan \beta = 2$ and $m_0 = 200$ GeV. Moderate values of $|\lambda v_s|$ are obtained.

In the nMSSM, the relic density of the lightest neutralino strongly depends on its mass [2, 3]. Although the dominant contribution to the annihilation cross section is schannel Z boson exchange, the coupling between the Z boson and \tilde{N}_1 is significantly small. This is because the lightest neutralino, \tilde{N}_1 is mainly composed of the fermionic component



Figure 4. The dependence of the Higgs mass on the SUSY breaking parameter is shown. In the left figure we set d = 2.5, and in the right figure we set $D_Y = 6$. Other parameters are chosen as $\lambda = 0.69$, tan $\beta = 2.0$ and $m_0 = 200$ GeV in both figures.

of the nMSSM gauge singlet, \hat{S} . The resonant effect near the Z pole mass is important for the sufficient annihilation of the lightest neutralino.

The neutralino mass matrix is

$$\left(\tilde{B} \ \tilde{W} \ \tilde{H}_{1}^{0} \ \tilde{H}_{2}^{0} \ \tilde{S}\right) \begin{pmatrix} m_{\lambda_{1}} & 0 & -c_{\beta}s_{w}M_{Z} & s_{\beta}s_{w}M_{Z} & 0\\ 0 & m_{\lambda_{2}} & c_{\beta}c_{w}M_{Z} & -s_{\beta}c_{w}M_{Z} & 0\\ -c_{\beta}c_{w}M_{Z} & c_{\beta}c_{w}M_{Z} & 0 & \mu_{\text{eff}} & \lambda v_{2}\\ s_{\beta}s_{w}M_{Z} & -s_{\beta}c_{w}M_{Z} & \mu_{\text{eff}} & 0 & \lambda v_{1}\\ 0 & 0 & \lambda v_{2} & \lambda v_{1} & 0 \end{pmatrix} \begin{pmatrix} \tilde{B} \\ \tilde{W} \\ \tilde{H}_{1}^{0} \\ \tilde{H}_{2}^{0} \\ \tilde{S} \end{pmatrix}, \quad (4.10)$$

where $s_{\beta} = \sin \beta$, $c_{\beta} = \cos \beta$ and $s_w = \sin \theta_W$. \tilde{B} , \tilde{W} , $\tilde{H}^0_{1,2}$ and \tilde{S} denote the bino, wino, higgsino and singlino respectively. The mass of the lightest neutralino, m_{χ_1} , becomes heavier as $|\mu_{\text{eff}}|$ decrease. This is because the mixing of the higgsinos becomes small as one can see from eq. (4.10). Since $|\mu_{\text{eff}}|$ is a decreasing function of D_Y , a larger D_Y leads to a larger m_{χ_1} . The dependence of the lightest neutralino mass on D_Y is shown in figure 5. On the other hand, a larger d leads to a smaller m_{χ_1} .

Figure 5 shows m_{χ} and the relic density of the neutralino, $\Omega_{\chi}h^2$. m_{χ} and $\Omega_{\chi}h^2$ are calculated with NMSSMTools and micrOMEGAs [22, 23]. When m_{χ} is large and close to m_Z , $\Omega_{\chi}h^2$ is small. The observed relic density of dark matter is given by [24, 25]

$$0.094 < \Omega_{\rm CDM} h^2 < 0.136. \tag{4.11}$$

This condition is satisfied with $m_{\chi} \simeq 35$ GeV. With such light dark matter, there are strong limits for the spin-independent WIMP-nucleon scattering cross section from CDMS [26] and XENON10 [27]. The strongest limit for the cross section is for it to be less than 5×10^{-44} cm² for $m_{\chi_1} \simeq 30$ GeV.

The spin-independent WIMP-nucleon elastic scattering cross section is written as

$$\sigma^{\rm SI} = \frac{4m_{\chi}^2 m_{\rm nucleus}^2}{\pi (m_{\chi} + m_{\rm nucleus})^2} \left[Zf_p + (A - Z)f_n \right]^2 .$$
(4.12)

 $f_{p,n}$ is the coupling between the WIMP and a nucleon given by [28]

$$f_{p,n} = \sum_{q=u,d,s} f_{T_q}^{(p,n)} a_q \frac{m_{p,n}}{m_q} + \frac{2}{27} f_{T_G}^{(p,n)} \sum_{q=c,b,t} a_q \frac{m_{p,n}}{m_q}.$$
(4.13)



Figure 5. The mass and the relic density of the lightest neutralino are shown as functions of the SUSY breaking parameter D_Y . The other parameters are chosen as $m_0 = 200 \text{ GeV}, d = 2.25, \lambda = 0.69$ and $\tan \beta = 2.0$.

 a_q are the WIMP-quark couplings. We focus on the dark matter-proton scattering cross section in the following discussion. The parameter f_{T_q} is defined by

$$f_{T_q} \equiv m_q \left\langle N | \bar{q}q | N \right\rangle \equiv m_q B_q, \tag{4.14}$$

and $f_{TG} = 1 - \sum_{q=u,d,s} f_{T_q}$. f_{T_q} can be written as [29]

$$f_{T_u} = \frac{m_u B_u}{m_p} = \frac{2\sigma_{\pi N}}{m_p \left(1 + \frac{m_d}{m_u}\right) \left(1 + \frac{B_d}{B_u}\right)},$$

$$f_{T_d} = \frac{m_d B_d}{m_p} = \frac{2\sigma_{\pi N}}{m_p \left(1 + \frac{m_u}{m_d}\right) \left(1 + \frac{B_u}{B_d}\right)},$$

$$f_{T_s} = \frac{m_s B_s}{m_p} = \frac{y\left(\frac{m_s}{m_d}\right)\sigma_{\pi N}}{m_p \left(1 + \frac{m_u}{m_d}\right)},$$
(4.15)

where $\sigma_{\pi N}$ is the π -nucleon sigma term:

$$\sigma_{\pi N} = \frac{1}{2} \left(m_u + m_d \right) \left(B_u + B_d \right).$$
(4.16)

The phenomenological value of $\sigma_{\pi N}$ is $64 \pm 8 \text{ MeV}$ [29]. y denotes the ratio of the strange quark component in the nucleon, defined as

$$y = \frac{2B_s}{B_u + B_d} . \tag{4.17}$$

y can be determined by the relation,

$$\sigma_0 = \sigma_{\pi N} \left(1 - y \right) = \frac{1}{2} \left(m_u + m_d \right) \left(B_u + B_d - 2B_s \right).$$
(4.18)

 σ_0 can be evaluated from baryon mass spectra using chiral perturbation theory. From [30], $\sigma_0 = 36 \pm 7$ MeV. There is large ambiguity for y. When $(\sigma_{\pi N}, \sigma_0) = (64, 36)$ MeV, y = 0.44. On the other hand, according to a recent lattice calculation [31], y has a small value such as 0.03.

The ratios of the quark mass are taken from [32].

$$\frac{m_u}{m_d} = 0.553 \pm 0.043, \qquad \qquad \frac{m_d}{m_s} = 18.9 \pm 0.8.$$
 (4.19)

The ratios of the form factors are written as

$$\frac{B_d}{B_u} = \frac{2 + (z-1)y}{2z - (z-1)y},\tag{4.20}$$

where

$$z = \frac{B_u - B_s}{B_d - B_s}.\tag{4.21}$$

z can be calculated from the baryon mass, and its value is 1.49 [33]. We can now determine f_{T_q} from eqs. (4.15), (4.19) and (4.20). When y = 0.44,

 $f_{T_u} \approx 0.027, \qquad f_{T_d} \approx 0.039, \qquad f_{T_s} \approx 0.365, \qquad f_{TG} \approx 0.569,$ (4.22)

and when y = 0.03,

$$f_{T_u} \approx 0.029, \qquad f_{T_d} \approx 0.036, \qquad f_{T_s} \approx 0.025, \qquad f_{TG} \approx 0.91.$$
 (4.23)

In these two cases, f_{T_s} and f_{TG} are very different. This affects the spin-independent cross section of the WIMP-nucleon scattering significantly.

WIMP-quark couplings, a_q , consist of two parts. One part arises from squark s-channel exchange and the other arises from the t-channel exchange of the neutral Higgs. The couplings from squark exchange are given by [34]

$$a_{q_i}^{\tilde{q}} = -\frac{1}{2(\tilde{m}_{1i}^2 - m_{\chi}^2)} \operatorname{Re}\left[X_i Y_i^*\right] - \frac{1}{2(\tilde{m}_{2i}^2 - m_{\chi}^2)} \operatorname{Re}\left[W_i V_i^*\right], \qquad (4.24)$$

where

$$X_{i} = \eta_{11}^{*} \frac{gm_{q_{i}} N_{1,5-i}^{*}}{2M_{w}B_{i}} - \eta_{12}^{*} e_{i}g' N_{11}^{*},$$

$$Y_{i} = \eta_{11}^{*} \left(\frac{y_{i}}{2}g' N_{11} + gT_{3i} N_{12}\right) + \eta_{12}^{*} \frac{gm_{q_{i}} N_{1,5-i}}{2M_{w}B_{i}},$$

$$W_{i} = \eta_{21}^{*} \frac{gm_{q_{i}} N_{1,5-i}^{*}}{2M_{w}B_{i}} - \eta_{22}^{*} e_{i}g' N_{11}^{*},$$

$$Y_{i} = \eta_{21}^{*} \left(\frac{y_{i}}{2}g' N_{11} + gT_{3i} N_{12}\right) + \eta_{22}^{*} \frac{gm_{q_{i}} N_{1,5-i}}{2M_{w}B_{i}},$$
(4.25)

and i = 1 for an up-type quark and i = 2 for a down-type quark. \tilde{m}_{1i} and \tilde{m}_{2i} denote a light squark mass and a heavy squark mass respectively. η denotes a squark mixing such that

$$\tilde{q}_l = \eta_{l1}\tilde{q}_L + \eta_{l2}\tilde{q}_R. \tag{4.26}$$



Figure 6. The spin-independent cross sections, σ^{SI} are shown. σ^{SI} are calculated with $m_0 = 200 \text{ GeV}$, $\lambda = 0.69$ and $\tan \beta = 2.0$. The upper three lines are calculated with y = 0.44. y is evaluated with chiral perturbation. The lower three lines are calculated with y = 0.03, which is the result from a recent lattice calculation.

 y_i , T_{3i} and e_i denote the hypercharge, isospin and electric charge of the quarks respectively. $B_1 = \sin \beta$ and $B_2 = \cos \beta$.

The couplings from neutral Higgs exchange in the nMSSM are given by [35]

$$a_{q_i}^h = \sum_{a=1}^3 \frac{1}{m_{h_a^0}^2} C_Y{}_a^i \operatorname{Re}[C_H^a], \qquad (4.27)$$

where

$$C_{Y_{a}^{i}} = -\frac{gm_{q_{i}}}{4M_{w}B_{i}}S_{a,3-i},$$

$$C_{H}^{a} = \left(-gN_{12}^{*} + g'N_{11}^{*}\right)\left(S_{a1}N_{13}^{*} - S_{a2}N_{14}^{*}\right)$$

$$-\sqrt{2}\lambda\left[S_{a3}N_{13}^{*}N_{14}^{*} + N_{15}^{*}\left(S_{a2}N_{13}^{*} + S_{a1}N_{14}^{*}\right)\right].$$
(4.28)

 S_{ij} denotes Higgs mixing. One can write the mass eigenstate of the Higgs as

$$h_a^0 = S_{a1}h_d^0 + S_{a2}h_u^0 + S_{a3}h_s. ag{4.29}$$

When $\lambda = 0$, eq. (4.28) agrees with the couplings in the MSSM given in [34].

Figure 6 shows the spin-independent cross section as a function of d and D_Y . When y = 0.44, σ^{SI} is already excluded by the current experiments. On the other hand, when y = 0.03, σ^{SI} is smaller than the upper limit from XENON10 in many regions of the parameter space. In this case, σ^{SI} is large enough to be detected or be excluded by the next-generation experiments.

4.4 Mass spectrum

Here we present the sparticle mass spectrum, the relic density of the neutralino and the spin-independent cross section of dark matter-proton scattering. The mass spectra are calculated using NMSSMTools, and the relic densities are calculated using micrOMEGAs.

			m_0	N_f	d	D_Y	λ	$\tan \beta$	3	
		input p1	200	1	2.2	4.84	0.69	2		
		p2	200	1	3.5	8.74	0.69	2		
	$ \mu_{\rm eff} $	$m_{H_{1}^{0}}$	$m_{H_{2}^{0}}$	n	$n_{H_{3}^{0}}$	$m_{A_{1}^{0}}$	m	A_{2}^{0}	$m_{H^{\pm}}$	$m_{\chi^0_1}$
output p1	264.2	127.2	328.8	40	60.1	285.5	48	7.8	433.8	34.2
p2	270.6	127.5	313.5	59	90.0	286.6	60	1.6	582.7	34.3
	$m_{\chi^0_2}$	$m_{\chi^0_3}$	$m_{\chi^0_4}$	r	$n_{\chi^0_5}$	$m_{\tilde{g}}$	m	χ_1^{\pm}	$m_{\chi_2^{\pm}}$	$m_{\tilde{\nu}_L}$
p1	198.3	310.5	336.4	40	03.8	262.1	19	7.2	350.2	174.5
p2	237.1	317.0	417.9	4	59.2	185.1	23	7.7	428.9	408.9
	$m_{\tilde{\nu}_{\tau}}$	$m_{\tilde{e}_L}$	$m_{\tilde{e}_R}$	r	$n_{\tilde{\tau}_1}$	$m_{\tilde{\tau}_2}$	m	\tilde{u}_L	$m_{\tilde{u}_R}$	$m_{\tilde{t}_1}$
p1	174.4	184.5	651.7	18	84.3	651.6	123	31.7	991.9	829.5
p2	408.8	413.1	925.1	41	13.1	925.1	168	32.2	1347.1	1126.7
	$m_{\tilde{t}_2}$	$m_{ ilde{d}_L}$	$m_{\tilde{d}_R}$	r	$n_{\tilde{b}_1}$	$m_{\tilde{b}_2}$	Ω	h^2	$\sigma_p^{SI}({ m cm}^2)$	
p1	1177.4	1233.2	1169.1	11	66.9	1170.1	1 0.1	.11	3.3×10^{-44}	
p2	1605.4	1683.3	1573.3	15	73.1	1598.9	9 0.1	31	3.1×10^{-44}	

Table 3. Mass spectra.

In this scenario, the gluino is light in a wide range of the parameter space. This is because the contributions from gauge mediation and anomaly mediation cancel. From eq. (3.4), the gluino mass at the messenger scale is written as

$$m_{\tilde{g}} = -g_3^2 \left(3 - d\right) m_0, \qquad (4.30)$$

where $m_0 = F_{\phi}/(16\pi^2)$. Particularly in the region where d = 3, the gluino mass $m_{\tilde{g}}$ vanishes.

The results of the numerical calculation are presented in table 3. When the deflection parameter d changes, the overall scale of the soft breaking terms changes. However, once we impose the condition of the observed relic density, μ_{eff} is almost determined by m_{χ} . Therefore the mass spectrum of the Higgs does not change significantly. Two samples in table 3 satisfy the current experimental limits from LEPII and XENON10 and also explain the observed relic abundance of dark matter. The mass of the gluino is $m_{\tilde{g}} \sim 200$ GeV. The lightness of the gluino is the characteristic feature of this scenario.

5 Conclusions

We investigated the phenomenology of the nMSSM with a Fayet-Iliopoulos D-term in the positively deflected anomaly mediation scenario.

In the deflected anomaly mediation scenario, the messenger sector is introduced. We showed that the couplings between the nMSSM fields and the messenger sector fields are forbidden by the discrete symmetry, and therefore the phenomenology at the weak scale is not affected by the detail of the messenger sector. We evaluated the soft breaking terms at the messenger scale without assuming small Yukawa couplings, and showed that the contributions from Yukawa couplings are the same as those of anomaly mediation. The soft breaking parameters are determined by the deflection parameter d, the messenger scale and contributions from the Fayet-Iliopoulos D-term.

We also discussed the phenomenology of the nMSSM at the weak scale. We found that electroweak symmetry breaking is successful, and moderate values of μ_{eff} are obtained. The mass of the lightest Higgs is heavier than the LEP bound. We also obtained sparticle mass spectra, and interestingly, the gluino is light.

We showed that the lightest neutralino is a good candidate for dark matter. The relic density explains the observed abundance of dark matter. The spin-independent dark matter-proton scattering cross section satisfies the upper limit from XENON10 when we consider a small value of the strange quark content of the nucleon as indicated by a recent lattice calculation. The cross section is large enough to be detected or excluded by nextgeneration experiments of direct detection.

We consider this scenario phenomenologically viable. If the light gluino is discovered, it may imply that SUSY breaking is mediated by supergravity and messengers, and these two effects are comparable.

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A Explicit example of the positively-deflected anomaly mediation scenario in the nMSSM

In this appendix, we show that the positively deflected anomaly mediation scenario is achieved in the nMSSM.

Let us discuss the scalar potential of X. After rescaling the fields as $\hat{X}\phi \to \hat{X}$, the superpotential of \hat{X} is

$$W(\hat{X}) = \frac{m_X}{2}\hat{\phi}\hat{X}^2.$$
 (A.1)

We also consider tadpoles generated by the supergravity interaction.

$$\Delta W(\hat{X}) = \mu_X^2 \hat{X},\tag{A.2}$$

and

$$\Delta \mathcal{L} = c_X \mu_X^2 X + h.c. \tag{A.3}$$

The tadpoles are expected to be much smaller than the plank scale owing to $Z_{7R'}$. We assume that all couplings of positive mass dimension in the messenger sector are on the order of F_{ϕ} : $m_X \sim \mu_X \sim c_X \sim F_{\phi}$.

From eq. (A.1), (A.3), we obtain the scalar potential for X,

$$V(X) = \left|\mu_X^2 + m_X X\right|^2 - \left(c_X \mu_X^2 X - \frac{m_X}{2} F_{\phi} X^2 + h.c.\right)$$
(A.4)

We assume that CP is conserved for simplicity. The messenger scale, X, is determined by the stationary condition and

$$X = \frac{\mu_X^2 (c_X - m_X)}{m_X (m_X - F_{\phi})}.$$
 (A.5)

The minimum condition, $\frac{\partial^2 V}{\partial X^2} > 0$ leads to $m_X > F_{\phi}$; therefore, the fermionic partner of X is not the LSP. We now evaluate the deflection parameter:

$$dF_{\phi} \equiv \frac{F_X}{X} - F_{\phi} = \frac{m_X(m_X - F_{\phi}) - (m_X + F_{\phi})(m_X - c_X)}{m_X - c_X}.$$
 (A.6)

When c_X is close to m_X , d is positive and large.

B Soft breaking masses and trilinear couplings

In this appendix, we present derivations for the soft breaking mass terms and the scalar trilinear couplings. We show that at the messenger scale, Yukawa contributions to the soft breaking terms are the same as those for anomaly mediation.

We write the anomalous dimension and beta-functions as

$$\gamma_i \equiv \frac{d\ln Z_i}{d\ln \mu} = \frac{1}{16\pi^2} \sum_A c_A^i g_A^2, \tag{B.1}$$

$$\frac{dg_a}{d\ln\mu} = -\frac{b_a}{16\pi^2}g_a^3,$$
(B.2)

$$\frac{dy}{d\ln\mu} = \frac{1}{16\pi^2} \left(d_a g_a^2 + d_y y^2 \right) y,$$
 (B.3)

where $g_A = \{g_1, g_2, g_3, y\}$. The coefficients of the gauge-coupling beta-function are $b_a - N_f$ above the messenger scale.

We obtain the wavefunction renormalization constant Z_i by integrating out the anomalous dimension γ_i in eq. (B.1) from μ to Λ .

$$\ln Z_i(s,t) = \ln Z_i(\Lambda) - \frac{1}{16\pi^2} \sum_A c_A^i \left(\int_s^t g_A^2(s',t) ds' + \int_t^{\ln\Lambda} g_A^2(t') dt' \right), \quad (B.4)$$

where $s = \ln \mu$ and $t = \ln |X|$. To obtain soft breaking terms, we need to differentiate $\ln Z_i(s,t)$ with respect to s and t. For this purpose, we define

$$I \equiv \int_{s}^{t} f(s',t)ds' + \int_{t}^{\ln\Lambda} f(t')dt', \qquad (B.5)$$

and differentiate this expression with respect to s and t:

$$\begin{aligned} \frac{\partial I}{\partial s} &= -f(s,t) \,,\\ \frac{\partial I}{\partial t} &= \int_{s}^{t} \frac{\partial f(s',t)}{\partial t} ds' + f(s,t)|_{s=t} - f(t) \,,\\ \frac{\partial^{2} I}{\partial s^{2}} &= -\frac{\partial f(s,t)}{\partial s} \,,\\ \frac{\partial I}{\partial s \partial t} &= -\frac{\partial f(s,t)}{\partial t} \,,\\ \frac{\partial I}{\partial t^{2}} &= \int_{s}^{t} \frac{\partial^{2} f(s',t)}{\partial t^{2}} ds' + \frac{\partial f(s,t)}{\partial s}\Big|_{s=t} + 2\frac{\partial f(s,t)}{\partial t}\Big|_{s=t} - \frac{\partial f(t)}{\partial t} \,. \end{aligned}$$
(B.6)

Using these formula, we obtain

$$\frac{\partial \ln Z_i(s,t)}{\partial s}\Big|_{s=t} = \sum_A \frac{c_A^i g_A^2}{16\pi^2},$$

$$\frac{\partial \ln Z_i(s,t)}{\partial t}\Big|_{s=t} = 0,$$

$$\frac{\partial^2 \ln Z_i(s,t)}{\partial s^2}\Big|_{s=t} = \sum_A \frac{c_A^i g_A(t)}{8\pi^2} \frac{\partial g_A(s,t)}{\partial s}\Big|_{s=t},$$

$$\frac{\partial^2 \ln Z_i(s,t)}{\partial s\partial t}\Big|_{s=t} = \sum_A \frac{c_A^i g_A(t)}{8\pi^2} \frac{\partial g_A(s,t)}{\partial t}\Big|_{s=t},$$

$$\frac{\partial^2 \ln Z_i(s,t)}{\partial t^2}\Big|_{s=t} = -\sum_A \frac{c_A^i g_A(t)}{8\pi^2} \left[\frac{\partial g_A(s,t)}{\partial s}, +2\frac{\partial g_A(s,t)}{\partial t} - \frac{\partial g_A(t)}{\partial t}\right]_{s=t}.$$
(B.7)

From eqs. (3.4) and (B.7), the scalar trilinear coupling is

$$a_{ijk}(t) = -\frac{F_{\phi}}{2} \left[\gamma_i(t) + \gamma_j(t) + \gamma_k(t)\right] y_{ijk}.$$
(B.8)

Next we derive soft breaking masses. For the gauge coupling

$$\frac{\partial g_a(s,t)}{\partial s} = -\frac{b_a}{16\pi^2} g_a^3,$$

$$\frac{\partial g_a(s,t)}{\partial t} = \frac{N_f}{16\pi^2} g_a^3,$$

$$\frac{\partial g_a(t)}{\partial t} = -\frac{b_a - N_f}{16\pi^2} g_a^3.$$
(B.9)

Therefore, from eqs. (3.4), (B.7) and (B.9), the soft breaking mass is

$$\tilde{m}_i^2(t) = \frac{|F_{\phi}|^2}{2(4\pi)^4} \sum_a c_a^i g_a^4(m) \left[b_a + N_f d(d+2) \right], \tag{B.10}$$

which agrees with [12, 14]. For Yukawa couplings,

$$y(s,t) = y(\Lambda) - \frac{1}{16\pi^2} \int_s^t ds' \left[d_a g_a^2(s',t) + d_y y^2(s',t) \right] y(s',t) - \frac{1}{16\pi^2} \int_t^{\ln\Lambda} dt' \left[d_a g_a^2(t') + d_y y^2(t') \right] y(t').$$
(B.11)

Using eq. (B.6), we obtain

$$\frac{\partial y(s,t)}{\partial s}\Big|_{s=t} = \frac{y(t)}{16\pi^2} \left[d_a g_a^2(s,t) + d_y y^2(s,t) \right]_{s=t} = \beta_y(t),$$

$$\frac{\partial y(s,t)}{\partial t}\Big|_{s=t} = 0,$$

$$\frac{\partial y(t)}{\partial t} = \beta_y(t).$$
(B.12)

Therefore Yukawa contributions to the soft breaking mass at the messenger scale are

$$\delta \tilde{m}_i^2 = -\frac{|F_\phi|^2}{4} \frac{c_y^i}{8\pi^2} y(t) \beta_y(t), \tag{B.13}$$

which are the same as those in anomaly mediation.

C Anomalous dimensions

One-loop anomalous dimensions of S, H_u and H_d are

$$\gamma_{s} = \frac{1}{16\pi^{2}} (-4\lambda^{2}),$$

$$\gamma_{H_{u}} = \frac{1}{16\pi^{2}} \left(\frac{3}{5}g_{1}^{2} + 3g_{2}^{2} - 6y_{t}^{2} - 2\lambda^{2} \right),$$

$$\gamma_{H_{d}} = \frac{1}{16\pi^{2}} \left(\frac{3}{5}g_{1}^{2} + 3g_{2}^{2} - 6y_{b}^{2} - 2y_{\tau}^{2} - 2\lambda^{2} \right).$$
(C.1)

At the one-loop level, the anomalous dimensions of other fields are the same as those in the MSSM. β_{λ} is

$$\beta_{\lambda} = -\frac{\lambda}{2} \left(\gamma_S + \gamma_{H_u} + \gamma_{H_d} \right). \tag{C.2}$$

References

- S.A. Abel, S. Sarkar and P.L. White, On the cosmological domain wall problem for the minimally extended supersymmetric standard model, Nucl. Phys. B 454 (1995) 663 [hep-ph/9506359] [SPIRES].
- C. Panagiotakopoulos and K. Tamvakis, New minimal extension of MSSM, Phys. Lett. B 469 (1999) 145 [hep-ph/9908351] [SPIRES].
- C. Panagiotakopoulos and K. Tamvakis, Stabilized NMSSM without domain walls, Phys. Lett. B 446 (1999) 224 [hep-ph/9809475] [SPIRES].
- [4] C. Panagiotakopoulos and A. Pilaftsis, Higgs scalars in the minimal non-minimal supersymmetric standard model, Phys. Rev. D 63 (2001) 055003 [hep-ph/0008268]
 [SPIRES].
- [5] A. Menon, D.E. Morrissey and C.E.M. Wagner, *Electroweak baryogenesis and dark matter in the NMSSM*, *Phys. Rev.* D 70 (2004) 035005 [hep-ph/0404184] [SPIRES].

- [6] S.J. Huber, T. Konstandin, T. Prokopec and M.G. Schmidt, *Electroweak phase transition and baryogenesis in the NMSSM*, Nucl. Phys. B 757 (2006) 172 [hep-ph/0606298] [SPIRES].
- [7] L. Randall and R. Sundrum, Out of this world supersymmetry breaking, Nucl. Phys. B 557 (1999) 79 [hep-th/9810155] [SPIRES].
- [8] A. Pomarol and R. Rattazzi, Sparticle masses from the superconformal anomaly, JHEP 05 (1999) 013 [hep-ph/9903448] [SPIRES].
- G.F. Giudice, M.A. Luty, H. Murayama and R. Rattazzi, Gaugino mass without singlets, JHEP 12 (1998) 027 [hep-ph/9810442] [SPIRES].
- [10] M. Ibe, R. Kitano and H. Murayama, A viable supersymmetric model with UV insensitive anomaly mediation, Phys. Rev. D 71 (2005) 075003 [hep-ph/0412200] [SPIRES].
- [11] R. Kitano, G.D. Kribs and H. Murayama, *Electroweak symmetry breaking via UV insensitive anomaly mediation*, *Phys. Rev.* D 70 (2004) 035001 [hep-ph/0402215] [SPIRES].
- [12] A. Pomarol and R. Rattazzi, Sparticle masses from the superconformal anomaly, JHEP 05 (1999) 013 [hep-ph/9903448] [SPIRES].
- [13] R. Rattazzi, A. Strumia and J.D. Wells, Phenomenology of deflected anomaly mediation, Nucl. Phys. B 576 (2000) 3 [hep-ph/9912390] [SPIRES].
- [14] N. Okada, Positively-deflected anomaly mediation, Phys. Rev. D 65 (2002) 115009
 [hep-ph/0202219] [SPIRES].
- [15] A. Cesarini, F. Fucito and A. Lionetto, Deflected anomaly mediation and neutralino dark matter, Phys. Rev. D 75 (2007) 025026 [hep-ph/0611098] [SPIRES].
- [16] P.M. Ferreira, I. Jack and D.R.T. Jones, *The three-loop SSM β-functions*, *Phys. Lett.* B 387 (1996) 80 [hep-ph/9605440] [SPIRES].
- [17] U. Ellwanger, J.F. Gunion and C. Hugonie, NMHDECAY: a Fortran code for the Higgs masses, couplings and decay widths in the NMSSM, JHEP 02 (2005) 066 [hep-ph/0406215]
 [SPIRES].
- [18] U. Ellwanger and C. Hugonie, NMHDECAY 2.0: an Updated program for sparticle masses, Higgs masses, couplings and decay widths in the NMSSM, Comput. Phys. Commun. 175 (2006) 290 [hep-ph/0508022] [SPIRES].
- [19] TEVATRON ELECTROWEAK WORKING GROUP collaboration, Combination of CDF and D0 results on the mass of the top quark, arXiv:0903.2503 [SPIRES].
- [20] A. Dedes, C. Hugonie, S. Moretti and K. Tamvakis, *Phenomenology of a new minimal supersymmetric extension of the standard model*, *Phys. Rev.* D 63 (2001) 055009
 [hep-ph/0009125] [SPIRES].
- [21] S.R. Coleman and E.J. Weinberg, Radiative corrections as the origin of spontaneous symmetry breaking, Phys. Rev. D 7 (1973) 1888 [SPIRES].
- [22] G. Bélanger, F. Boudjema, C. Hugonie, A. Pukhov and A. Semenov, *Relic density of dark matter in the NMSSM*, JCAP 09 (2005) 001 [hep-ph/0505142] [SPIRES].
- [23] G. Bélanger, F. Boudjema, A. Pukhov and A. Semenov, MicrOMEGAs2.0: a program to calculate the relic density of dark matter in a generic model, Comput. Phys. Commun. 176 (2007) 367 [hep-ph/0607059] [SPIRES].

- [24] WMAP collaboration, D.N. Spergel et al., Wilkinson Microwave Anisotropy Probe (WMAP) three year results: implications for cosmology, Astrophys. J. Suppl. 170 (2007) 377 [astro-ph/0603449] [SPIRES].
- [25] J. Hamann, S. Hannestad, M.S. Sloth and Y.Y.Y. Wong, How robust are inflation model and dark matter constraints from cosmological data?, Phys. Rev. D 75 (2007) 023522
 [astro-ph/0611582] [SPIRES].
- [26] CDMS collaboration, Z. Ahmed et al., Search for weakly interacting massive particles with the first five-tower data from the cryogenic dark matter search at the Soudan Underground Laboratory, Phys. Rev. Lett. 102 (2009) 011301 [arXiv:0802.3530] [SPIRES].
- [27] XENON collaboration, J. Angle et al., First results from the XENON10 dark matter experiment at the Gran Sasso National Laboratory, Phys. Rev. Lett. 100 (2008) 021303
 [arXiv:0706.0039] [SPIRES].
- [28] G. Jungman, M. Kamionkowski and K. Griest, Supersymmetric dark matter, Phys. Rept. 267 (1996) 195 [hep-ph/9506380] [SPIRES].
- [29] J.R. Ellis, K.A. Olive, Y. Santoso and V.C. Spanos, Update on the direct detection of supersymmetric dark matter, Phys. Rev. D 71 (2005) 095007 [hep-ph/0502001] [SPIRES].
- [30] B. Borasoy and U.-G. Meissner, Chiral expansion of baryon masses and σ-terms, Annals Phys. 254 (1997) 192 [hep-ph/9607432] [SPIRES].
- [31] H. Ohki et al., Nucleon sigma term and strange quark content from lattice QCD with exact chiral symmetry, Phys. Rev. D 78 (2008) 054502 [arXiv:0806.4744] [SPIRES].
- [32] H. Leutwyler, Light quark effective theory, hep-ph/9609465 [SPIRES].
- [33] H.-Y. Cheng, Low-energy interactions of scalar and pseudoscalar Higgs bosons with baryons, Phys. Lett. B 219 (1989) 347 [SPIRES].
- [34] J.R. Ellis, A. Ferstl and K.A. Olive, Re-evaluation of the elastic scattering of supersymmetric dark matter, Phys. Lett. B 481 (2000) 304 [hep-ph/0001005] [SPIRES].
- [35] D.G. Cerdeno, C. Hugonie, D.E. Lopez-Fogliani, C. Muñoz and A.M. Teixeira, Theoretical predictions for the direct detection of neutralino dark matter in the NMSSM, JHEP 12 (2004) 048 [hep-ph/0408102] [SPIRES].